

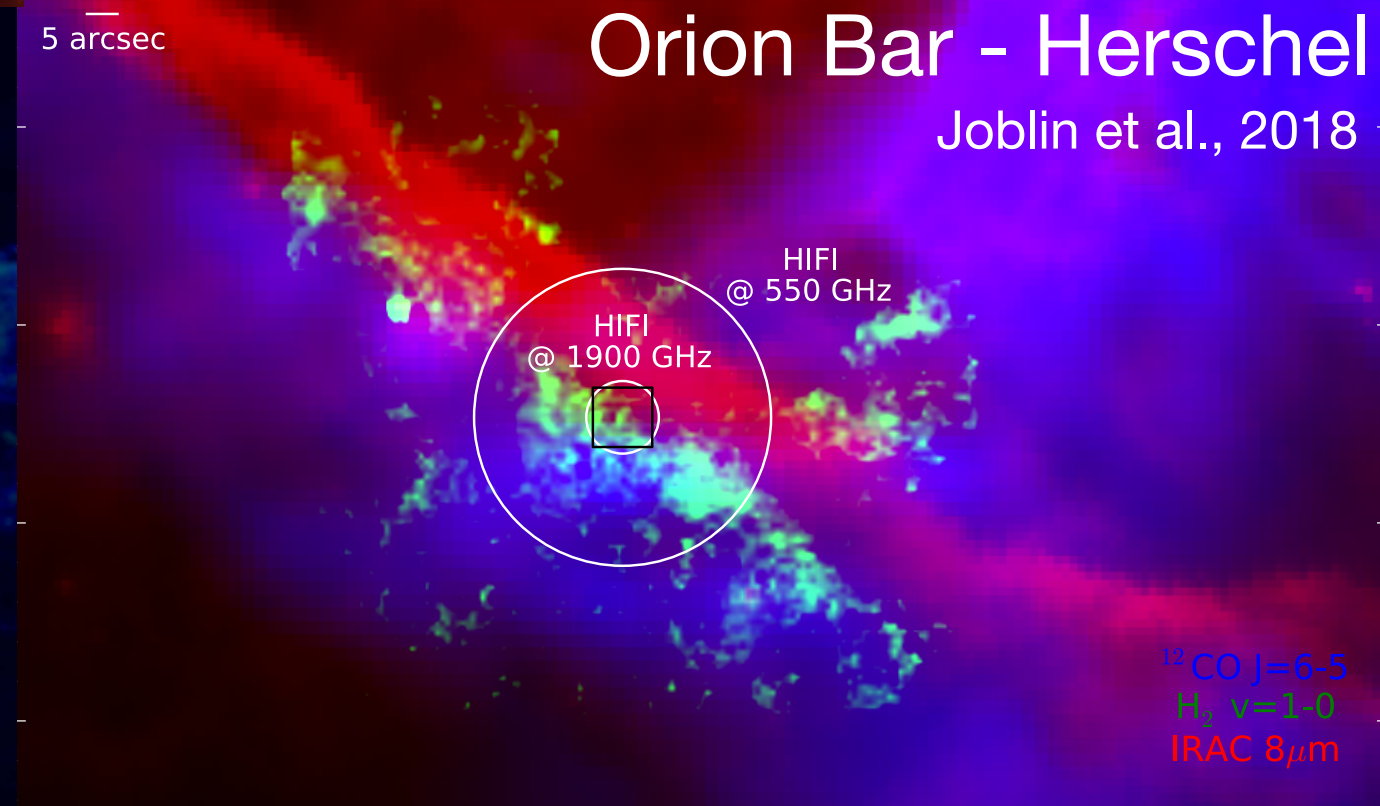
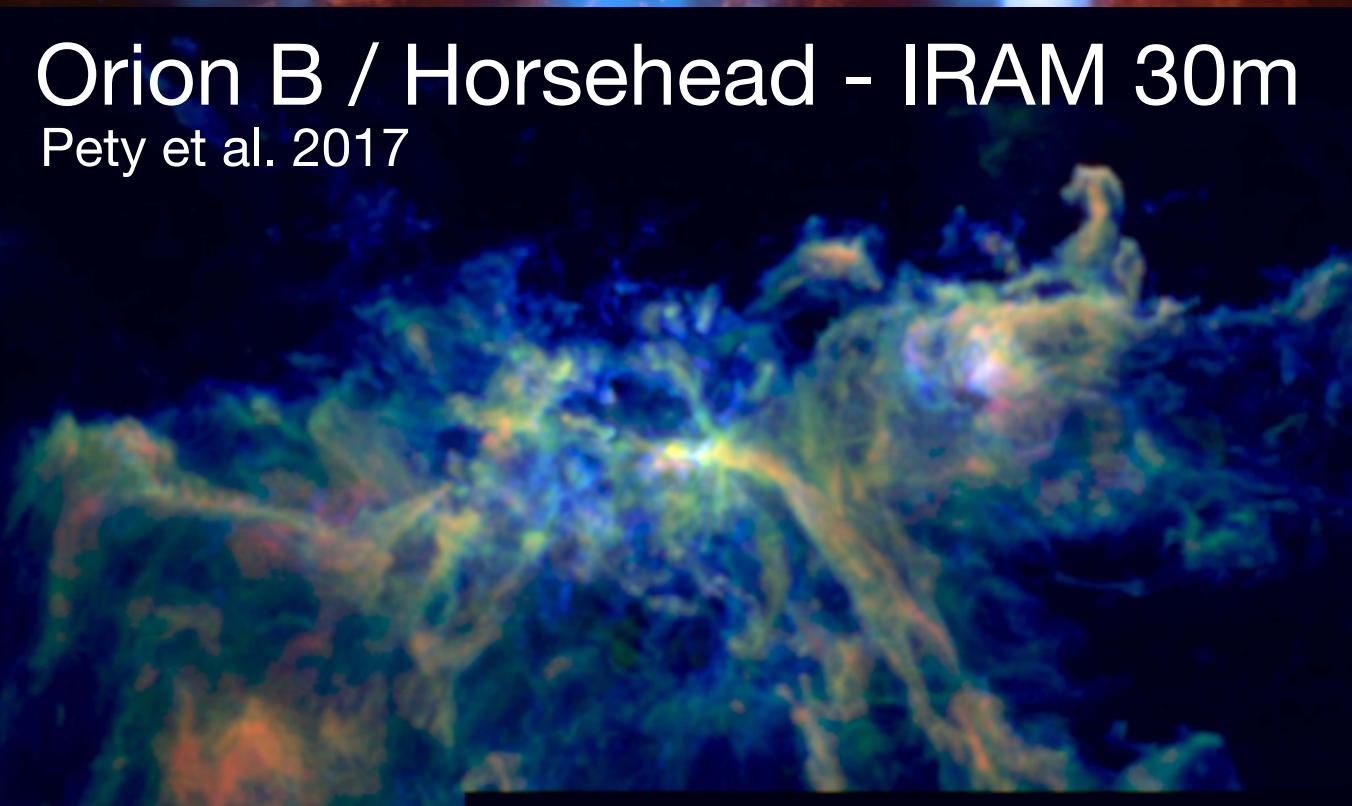
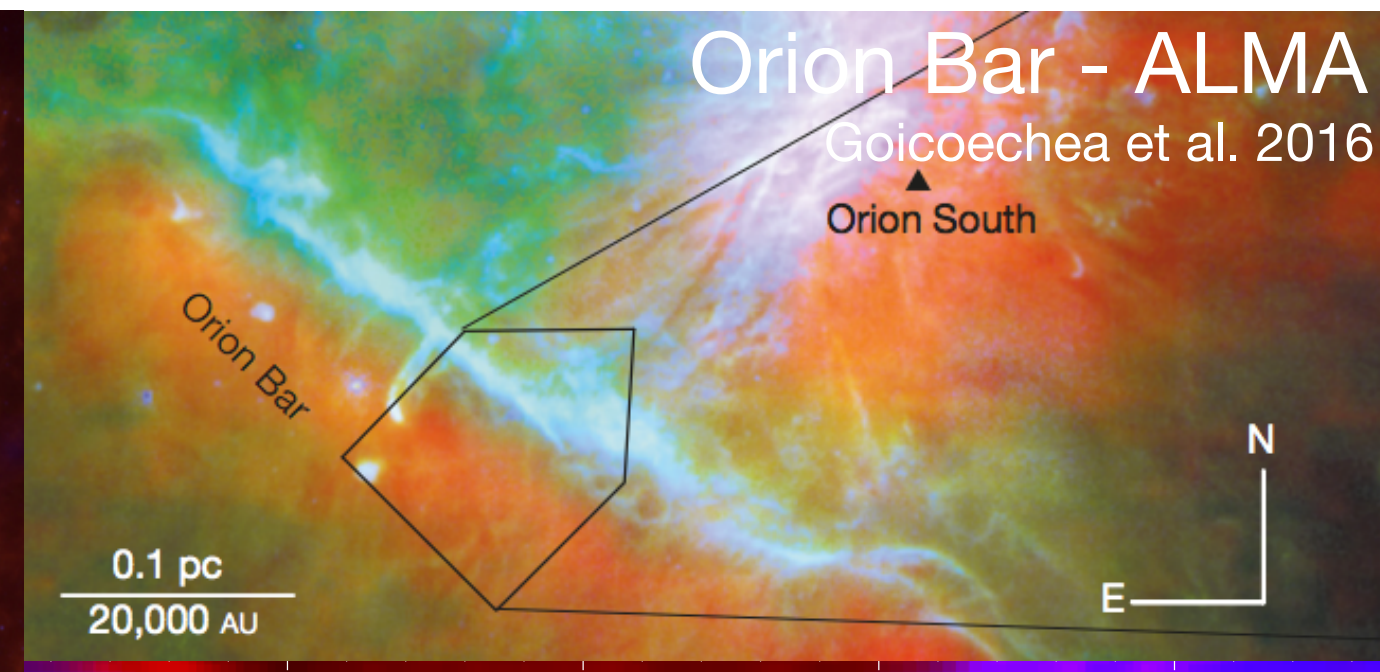
# New developments on the MIS & Jets platform

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Emeric Bron  
Franck Le Petit  
David Languignon  
Nicolas Moreau

One of the services of the ANO5 "Plateforme MIS & Jets"

Goal: Provide services to prepare and interpret observations in Galactic & extragalactic interstellar medium

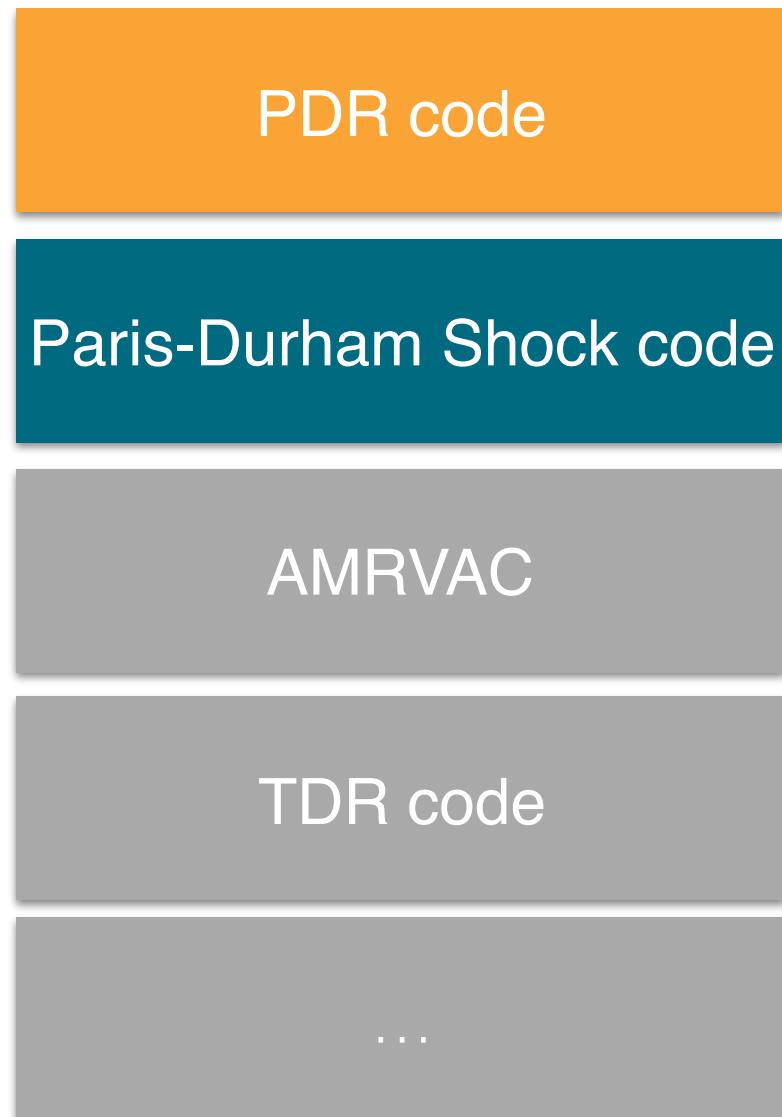




One of the services of the ANO5 "Plateforme MIS & Jets"

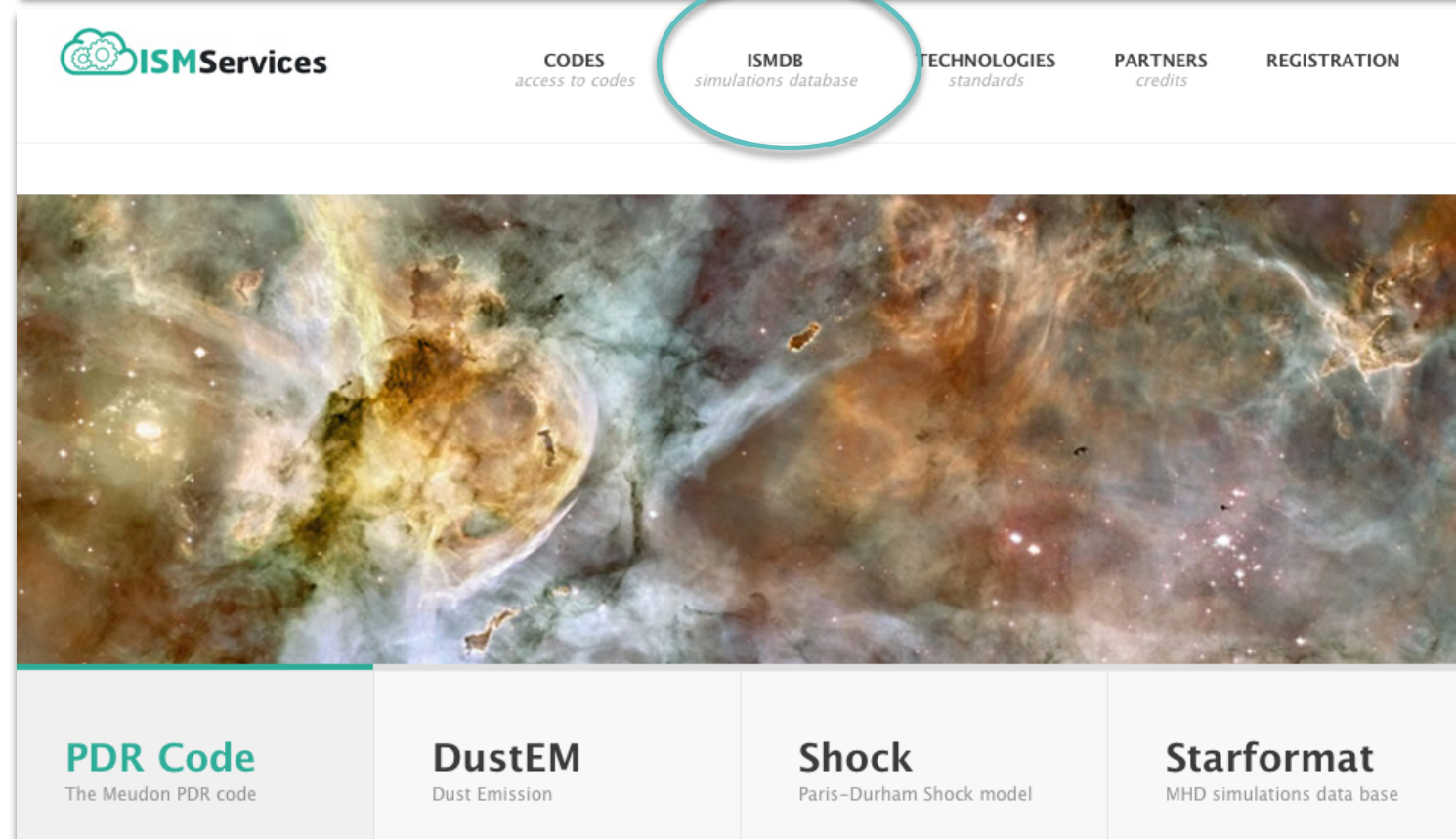
Goal: Provide services to prepare and interpret observations in Galactic & extragalactic interstellar medium

Services are based on reference state-of-the-art codes



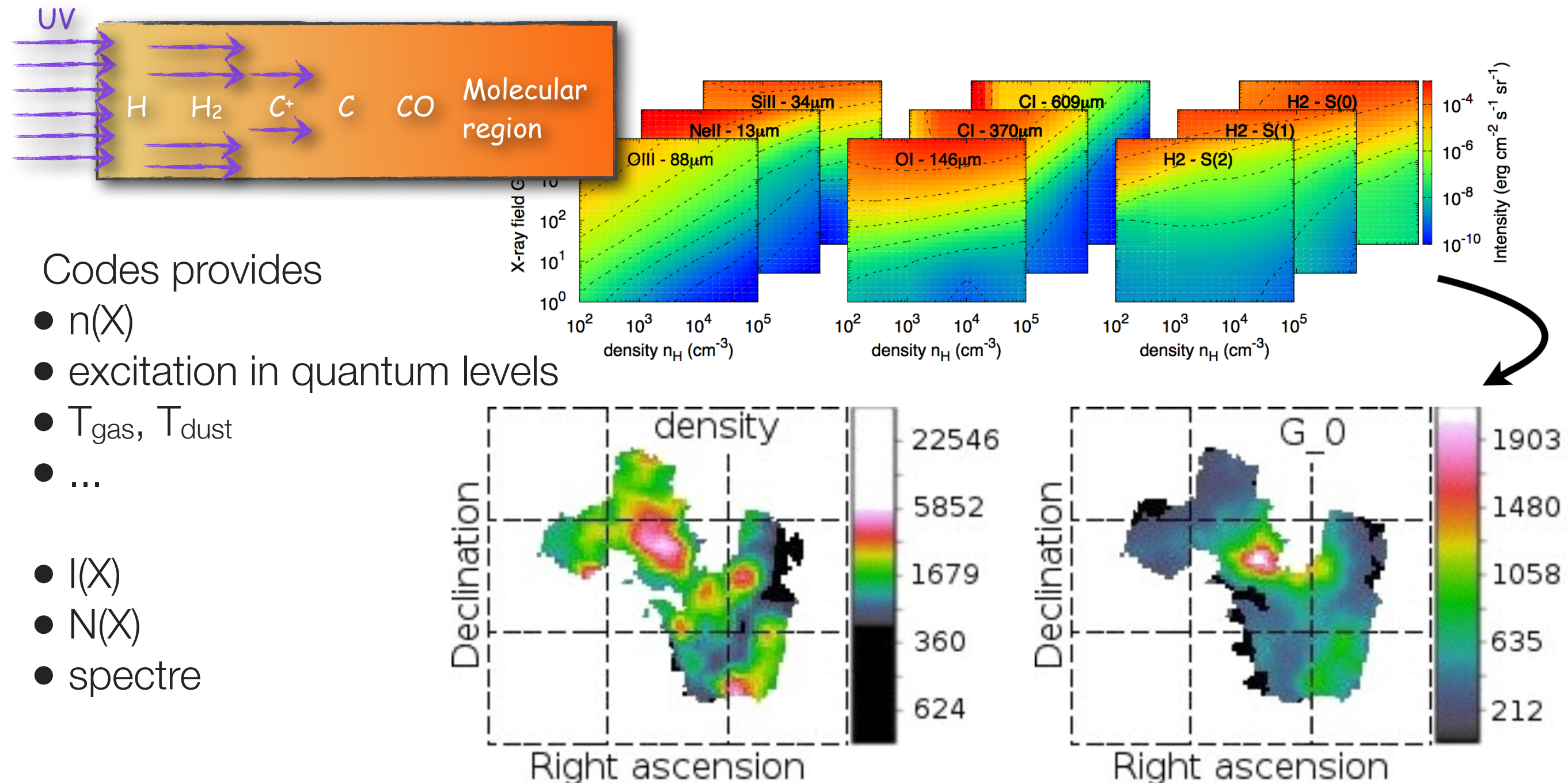
Several services are developed above the products of these codes

- Source codes & specific developments
- Online codes
- Tools to analyze results
  - Extractor & Chemistry Analyzer
- ISMDB



One of the services of the ANO5 "Plateforme MIS & Jets"

Goal: Provide services to prepare and interpret observations in Galactic & extragalactic interstellar medium

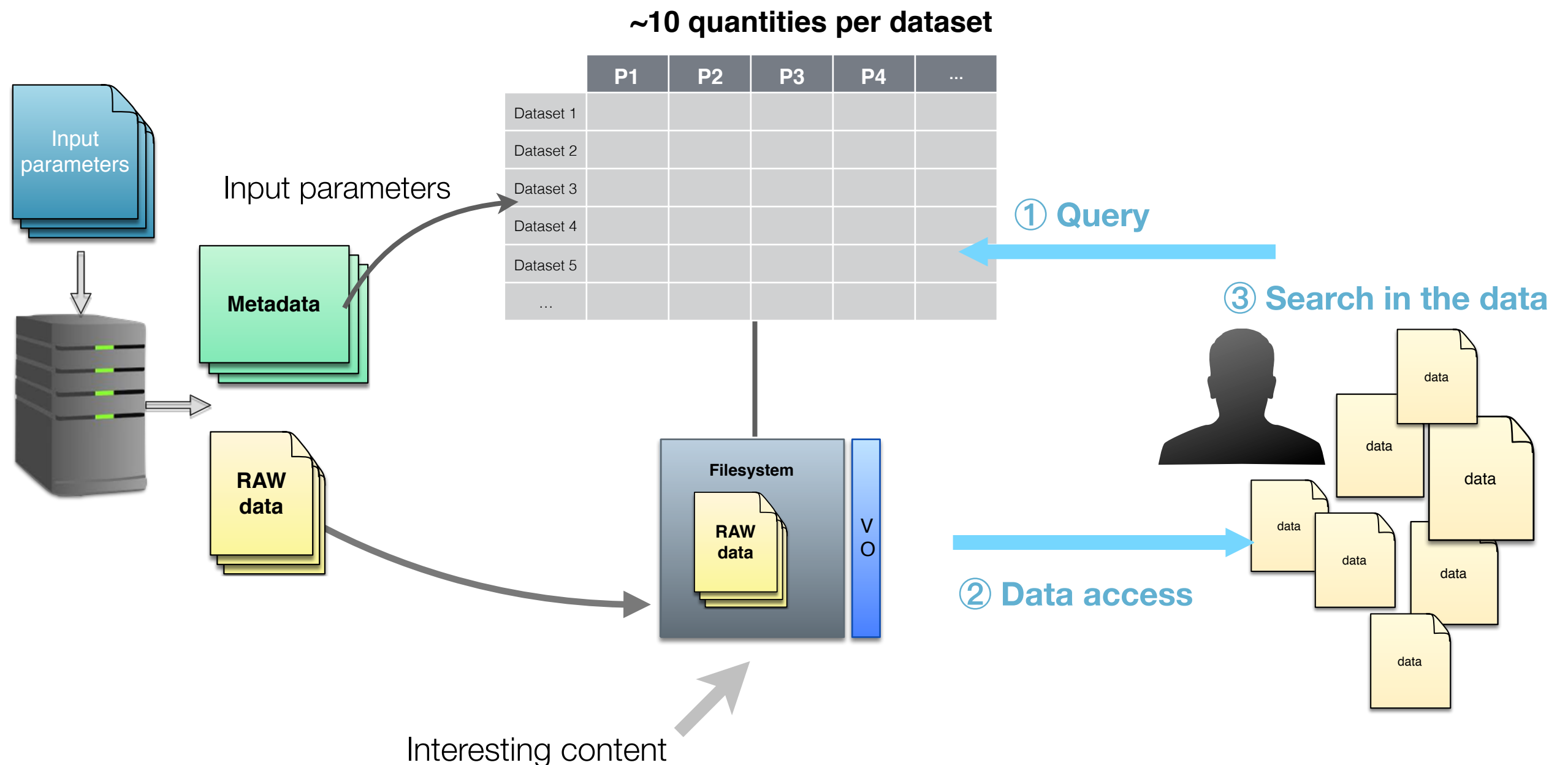




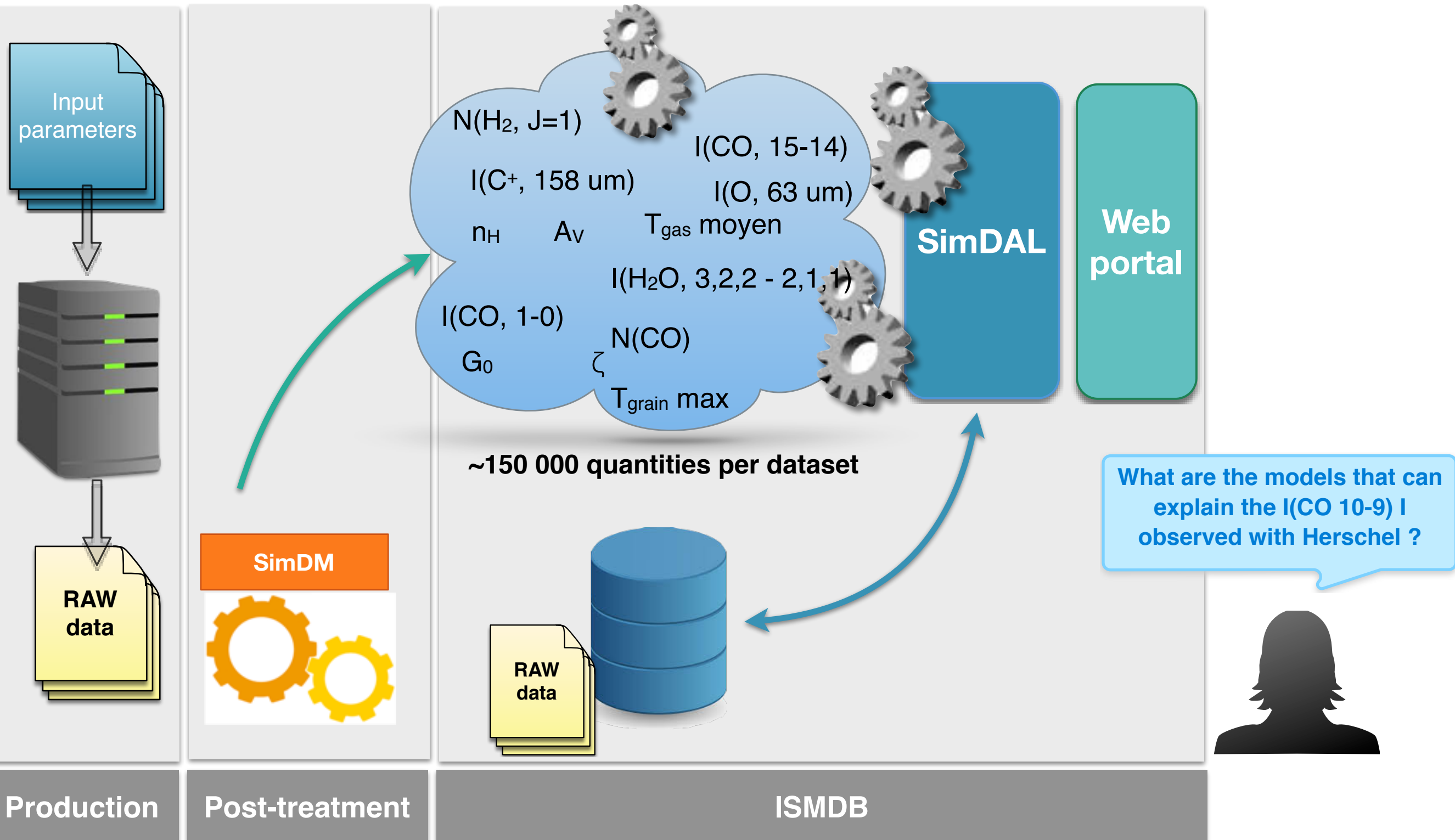
## ISMDB: InterStellar Medium DataBase

- not only a classical database to find pre-computed models
- but **also a tool that can *interpret* observations**

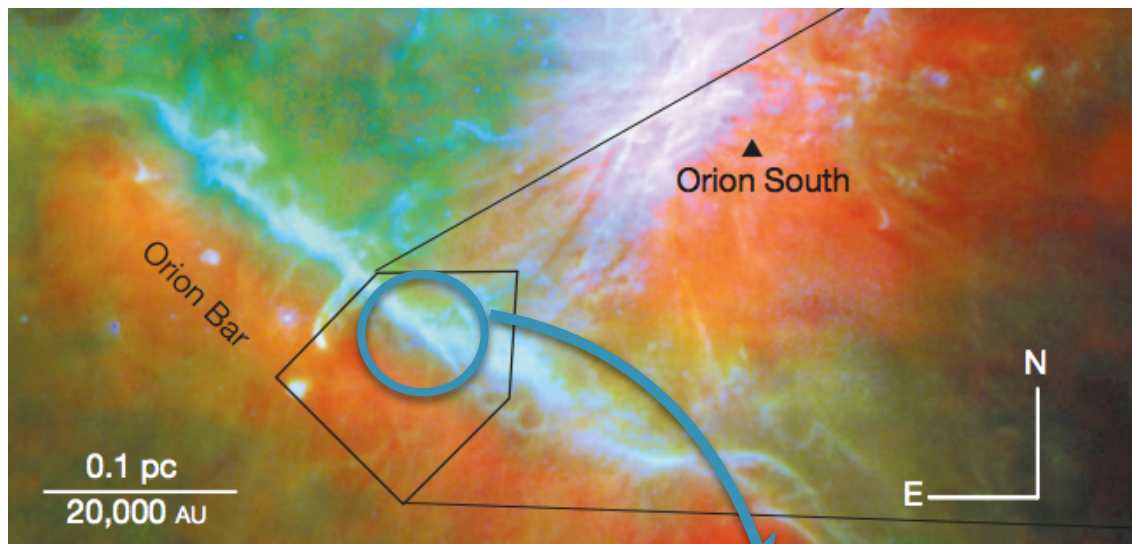
Standard databases:



- not only a classical database to find pre-computed models
- but **also a tool that can *interpret* observations**

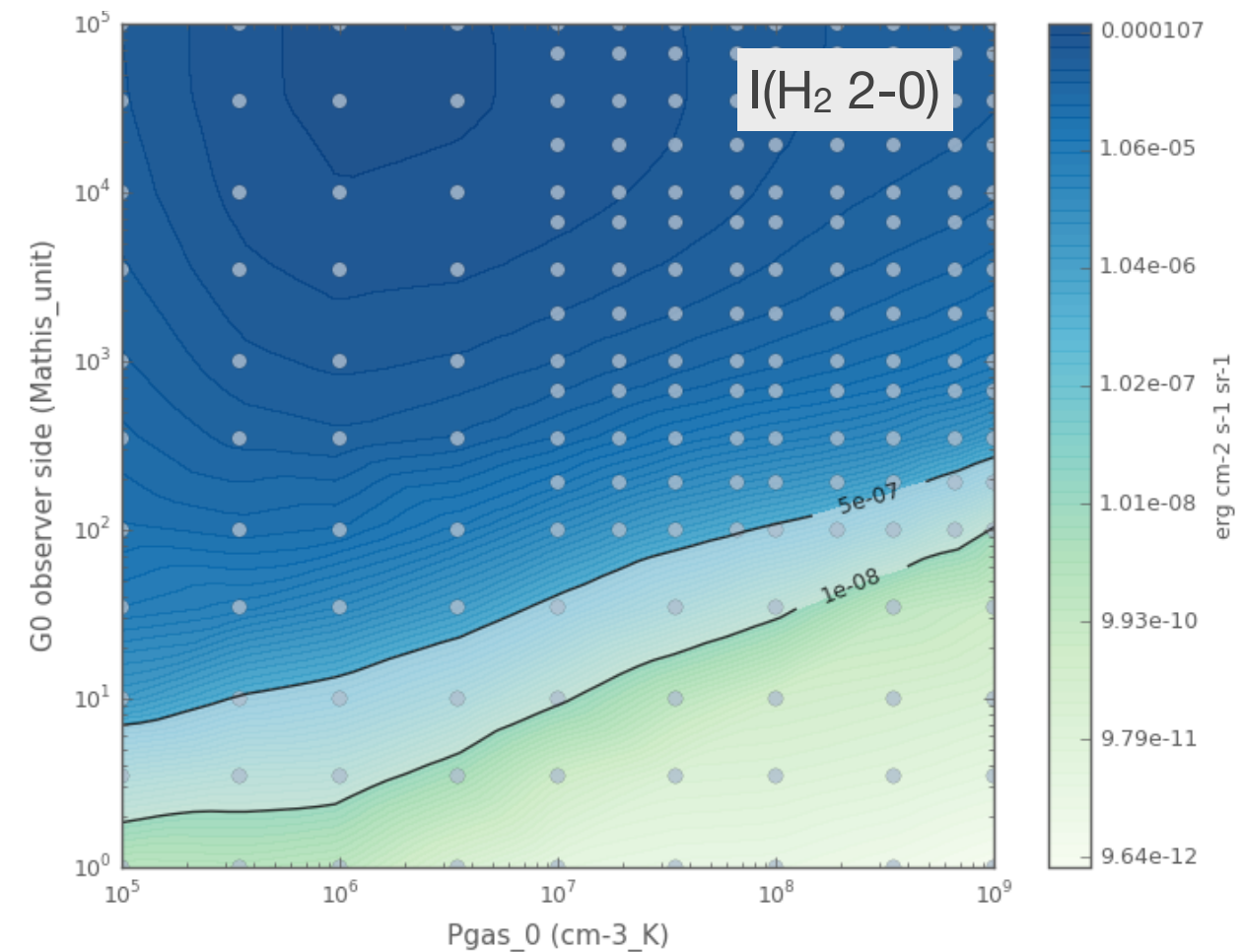
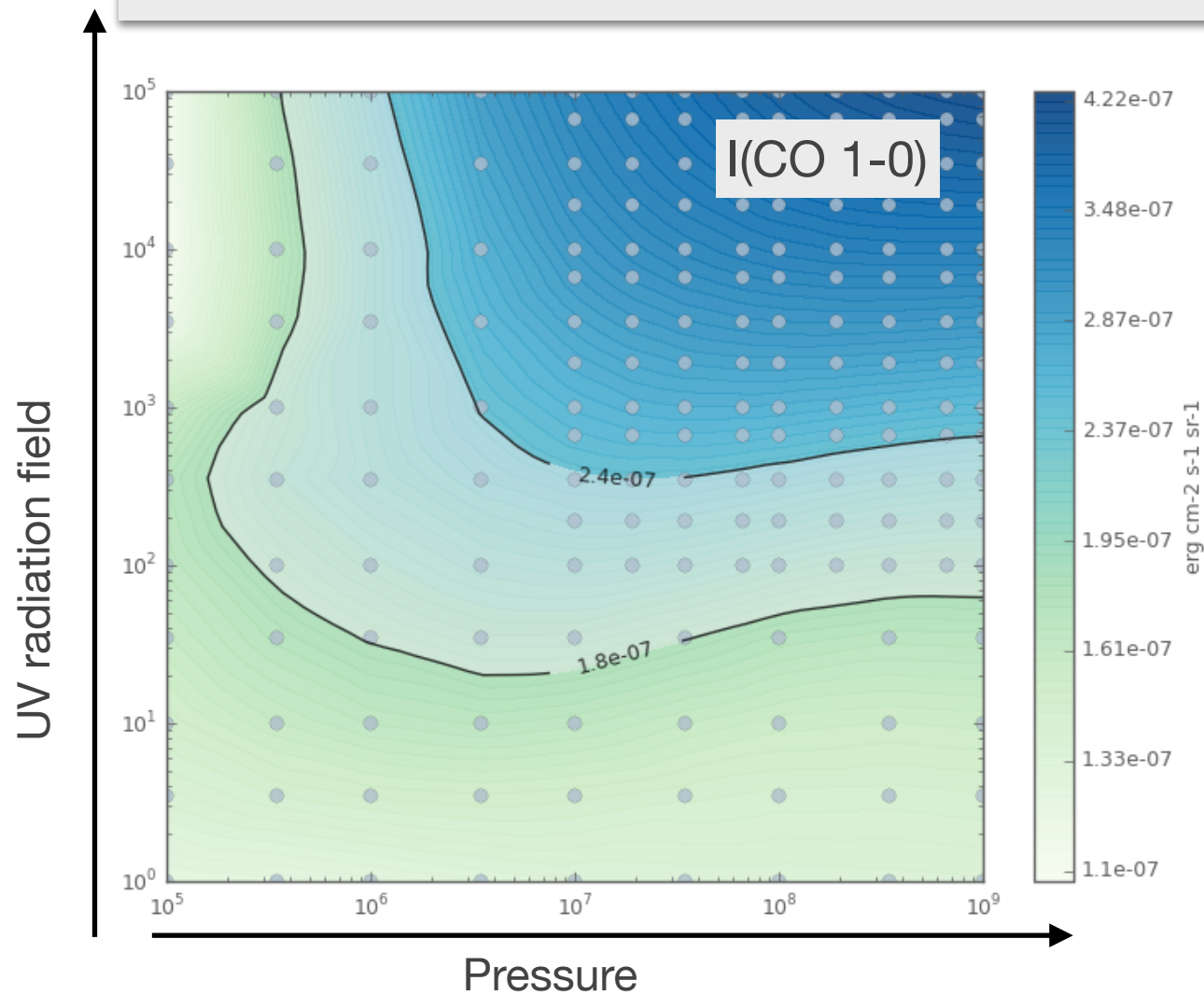
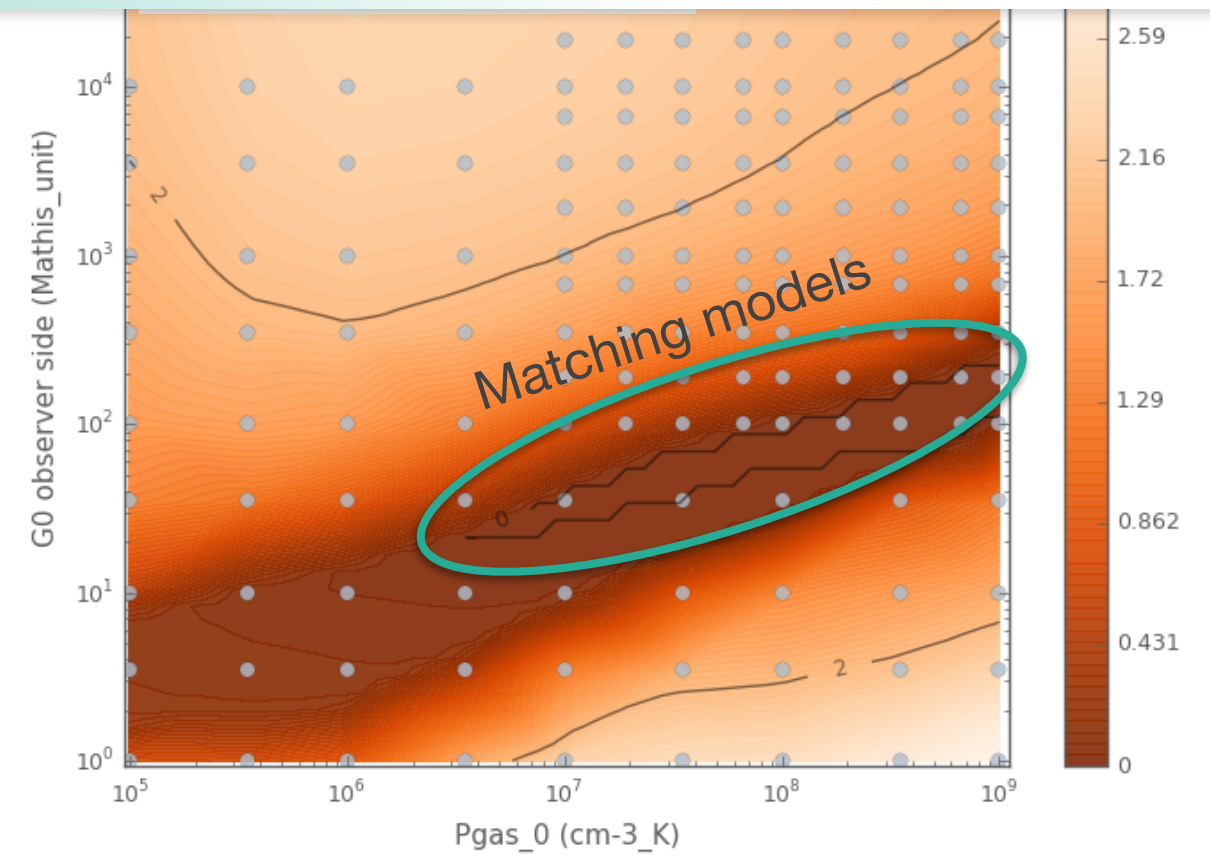






$$1.8 \cdot 10^{-7} < I(\text{CO } 1-0) < 2.4 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$

$$1.0 \cdot 10^{-8} < I(\text{H}_2 \text{ } 2-0) < 5.0 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$



## VLA archive Interface

**NRAO Science Data Archive : Advanced Search Tool**  
**Historical VLA, Jansky VLA, VLBA and GBT Data Products**

**Output Control Parameters :**

**Choose Query Return Type :**

- ☒ Download Archive Data Files
- ☐ VLA Observations Summary
- ☐ List of Observation Scans
- ☐ List of Projects

[Output Tbl Format](#)  [Sort Order Column 1](#)

[Max Output Tbl Rows](#)  [Sort Order Column 2](#)

**General Search Parameters :**

[Telescopes](#) ☒ All ☐ Jansky VLA ☐ Historical VLA ☐ VLBA ☐ GBT

[Project Code](#)   
JVLA: 12A-256

[Project Session](#)

[Dates From](#)

[Observer Name](#)

[Archive File ID](#)  (partial strings allowed)

[To](#)  (2010-06-21 14:20:30)

**Position Search :**

[Target Name](#)

[Search Type](#)

[Min. Exposure](#)  (secs)

[RA or Longitude](#)

[DEC or Latitude](#)

[Equinox](#)

[Search Radius](#)  (1d00'00" or 0.2d)

- OR - ☐ Check for automatic VLA field-of-view, freq. dependent.??

**Observing Configurations Search :**

[Telescope](#) ☒ All ☐ A ☐ AB ☐ BnA ☐ B ☐ BC ☐ CnB

[Config](#) ☐ C ☐ CD ☐ DnC ☐ D ☐ DA

[Sub\\_array](#) ☒ All ☐ 1 ☐ 2 ☐ 3 ☐ 4 ☐ 5

[Polarization](#)

[Data Type](#)

[Observing Bands](#) ☒ All ☐ 4 ☐ P ☐ L ☐ S ☐ C ☐ X ☐ U ☐ K ☐ Ka ☐ Q ☐ W

[Frequency Range](#)  (In MHz : 1665.401 - 1720.500)

[Enter Locked Project Access key :](#)

Unique keywords may be used to unlock proprietary data from individual observing projects. Contact the [NRAO Data Analysts](#) for project access keys.

23 paramètres de recherche  
Interface complexe

**ISMDB**  
**150 000 parameters !**



## ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models

2016.12.03

### 1 – search among two parameters

x  (cm<sup>-3</sup>\_K) ☒ log scale  
y  (Mathis\_unit) ☒ log scale

### 2 – fix all the other parameters

(mag)

### 3 – observational constraints

Use

```
"I(CO v=0,J=1->v=0,J=0 angle 00 deg)" > 1.8E-7  
"I(CO v=0,J=1->v=0,J=0 angle 00 deg)" < 2.4E-7  
"I(H2 v=0,J=2->v=0,J=0 angle 60 deg)" > 1E-8  
"I(H2 v=0,J=2->v=0,J=0 angle 60 deg)" < 5E-7
```

Search

### 3 – observational constraints

Use



Semantics interpreter

Standard activities continue:

- upgrade of codes:
  - new physics
  - new atomic, molecular & chemistry data
  - more detailed documentations
- continue formation of students (schools & master classes)
- participation in IVOA on semantics + transfer of competences on SimDAL

## Major new developments:

- ① Integration of the various services (e.g. extractor tools in web interface)
- ② Extend content of ISMDB (additional codes)
- ③ ISMDB evolutions : map interpretation
- ④ Links with other services



# ① Integration of the various services

- Integration of Analysis Tools on ISMDB web-page

- data extractor
- chemistry analyzer

→ Plot & extract any quantity online

→ Analyse results online

- New presentation of models:

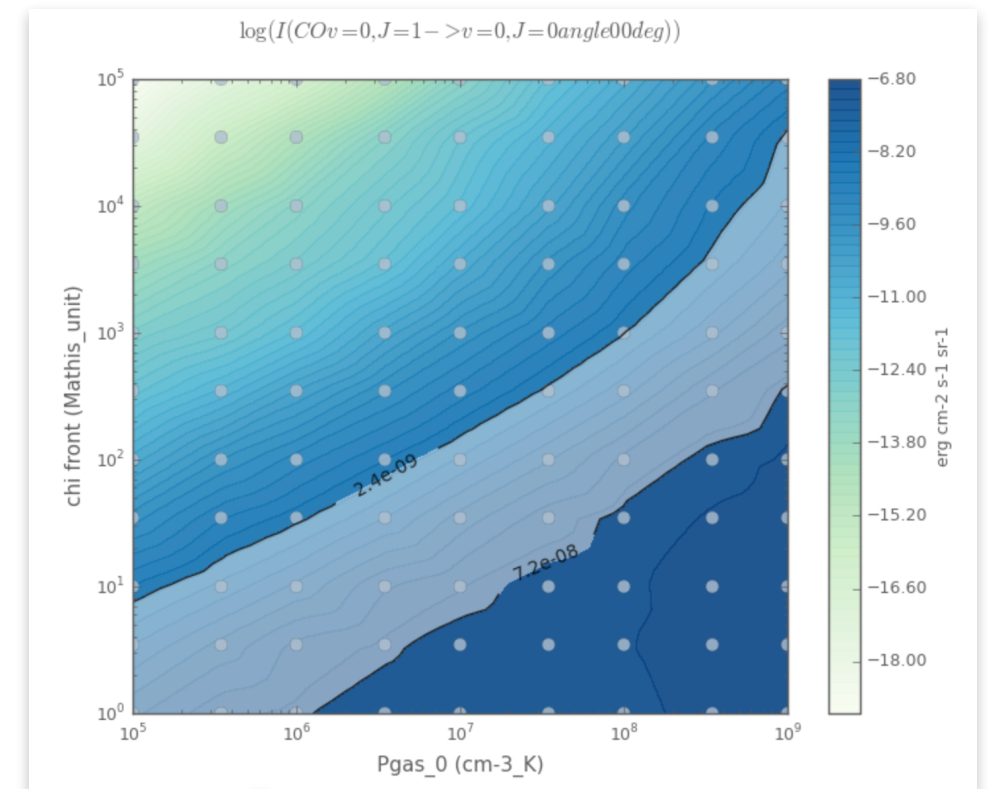
Each model presented as card with:

- **main results** (intensities, column densities)
- **pre-defined plots** of important quantities

- Download data file for all generated plots

- allows users to manipulate data as they wish

- Search by input parameters

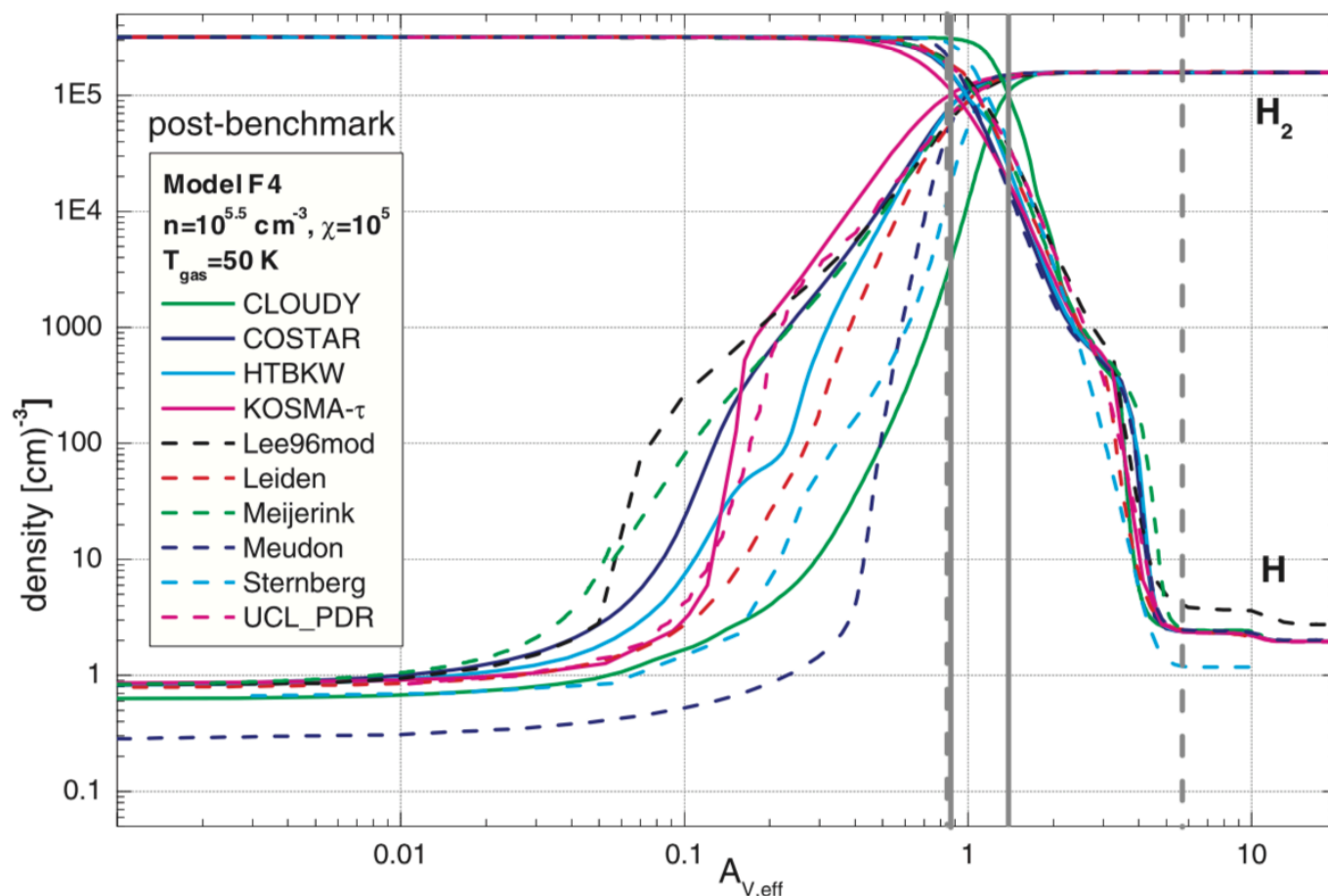


P	G0	I(CO 1-0)
...	...	...
...	...	...
...	...	...

## ② Extend scientific content of ISMDB

Several kinds of models are required to interpret observations in Galactic and extragalactic ISM:

- PDR
- shocks
- dust emission
- H II regions models
- radiative transfer
- astrochemistry models
- ...



Complex physics in these codes

- useful to have **several sources to compare results**

Röllig et al. (2006)

We have defined a **simple data format** to ingest models in ISMDB  
→ simple for data providers to integrate their models in ISMDB



## Presently in ISMDB:

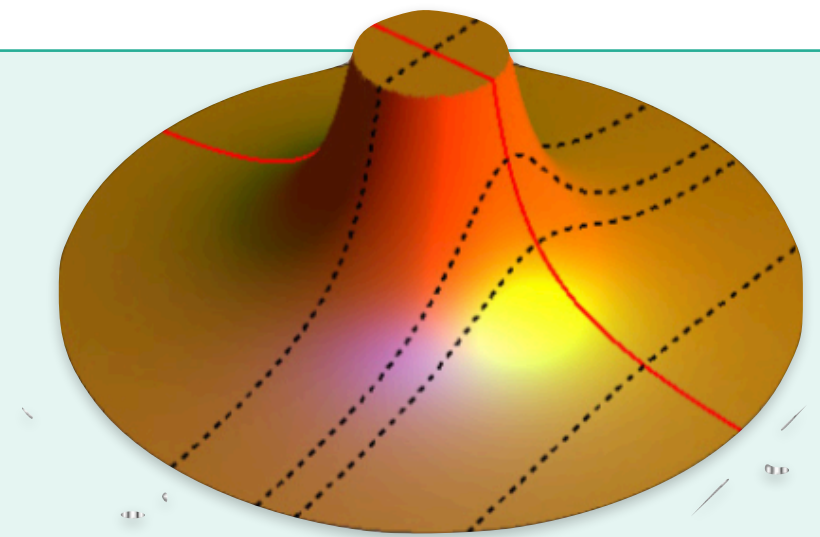
Two Meudon PDR code grids

- constant  $n_H$  and constant  $P$
- Galactic PDR conditions ( $Z = 1$ )
- $\sim 2000$  models

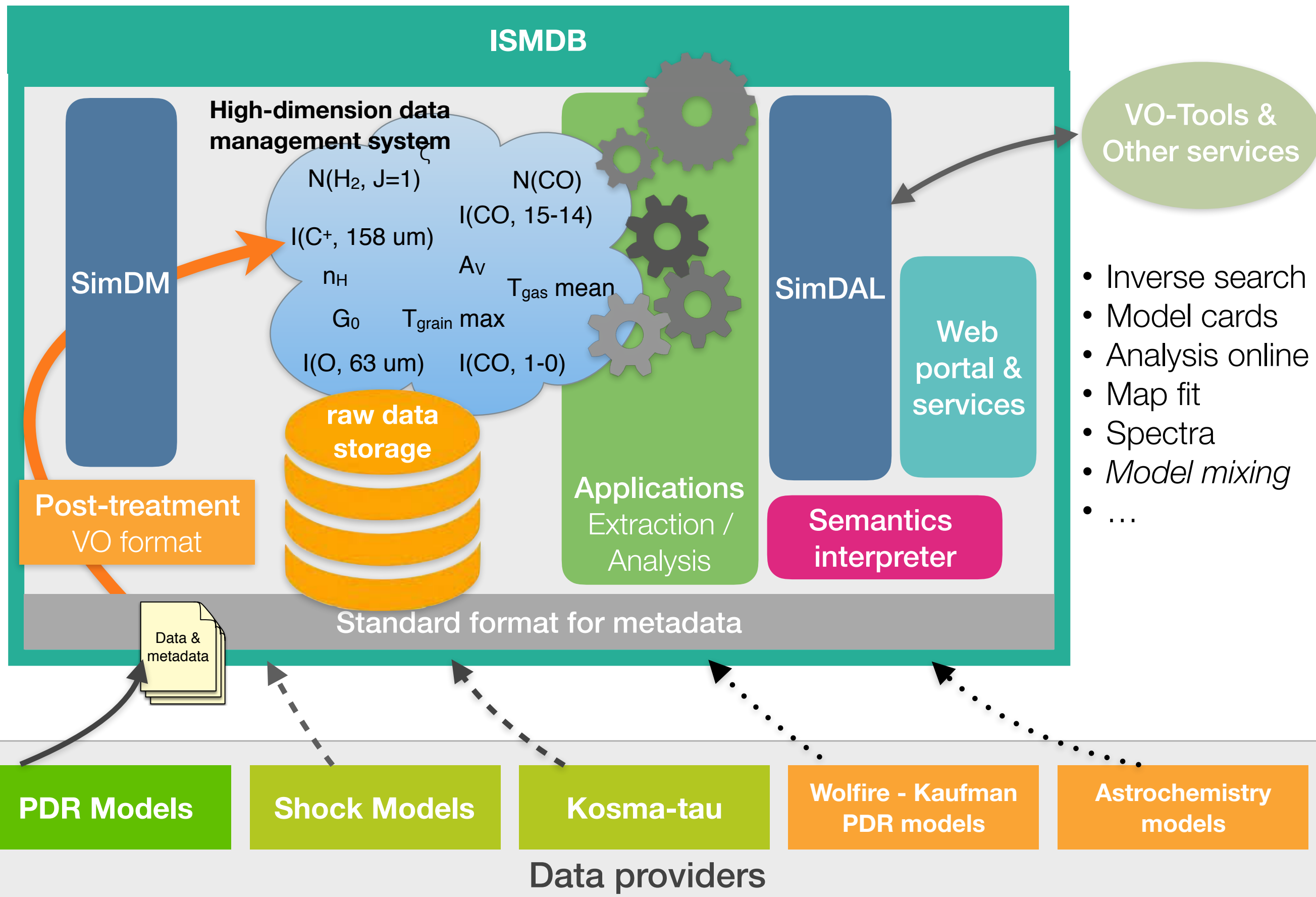
Extensions motivated by JWST & IRAM observations

In 2019, our goals:

- Meudon PDR code in **extragalactic conditions**
- **Paris-Durham shock** models (A. Gusdorf et al.)
- **Kosma- $\tau$  PDR models** (Cologne PDR code - Markus. Röllig)



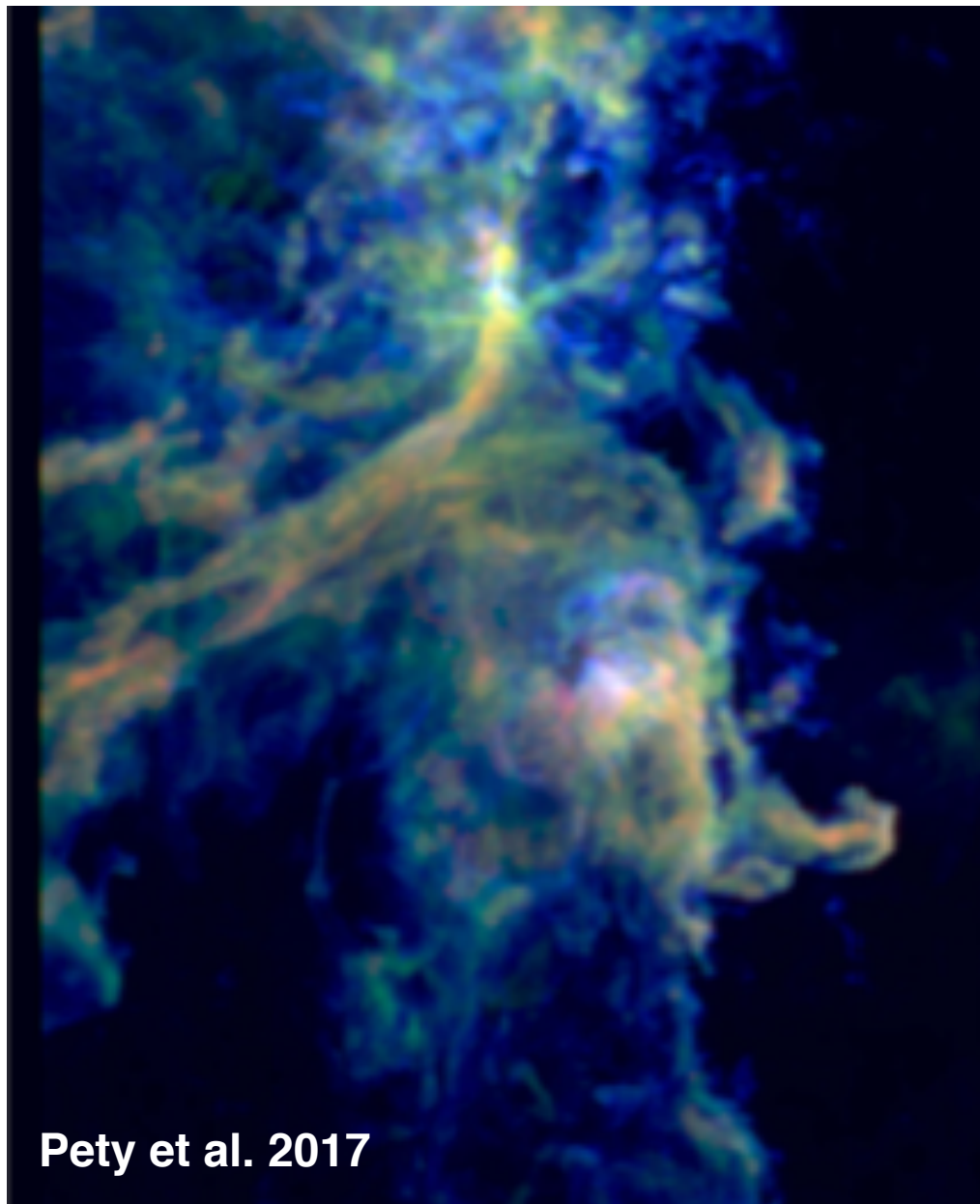
Profile of a clump  
in Kosma- $\tau$   
(Röllig et al. 2013)



### ③ New service: interpretation of maps

Motivation: more and more hyper-spectral observations  
need tools to interpret them

Instruments: JWST, IRAM, ALMA



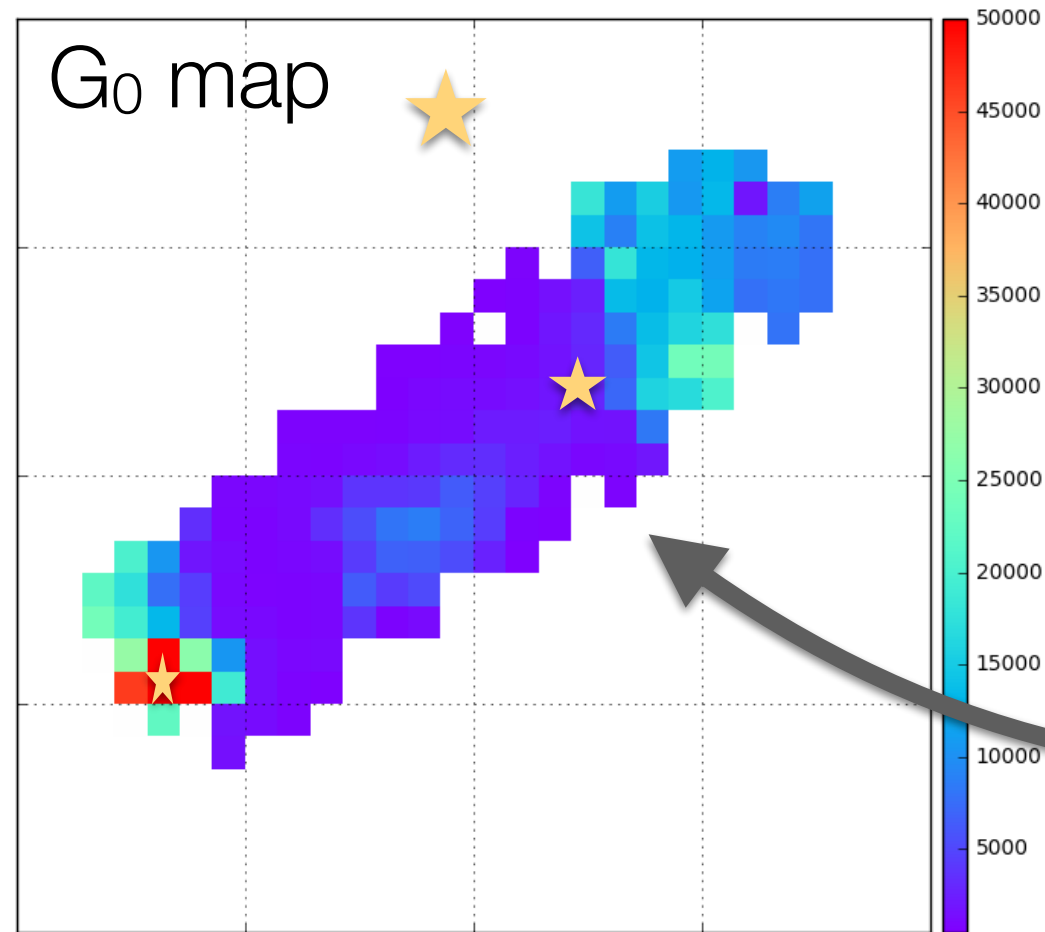
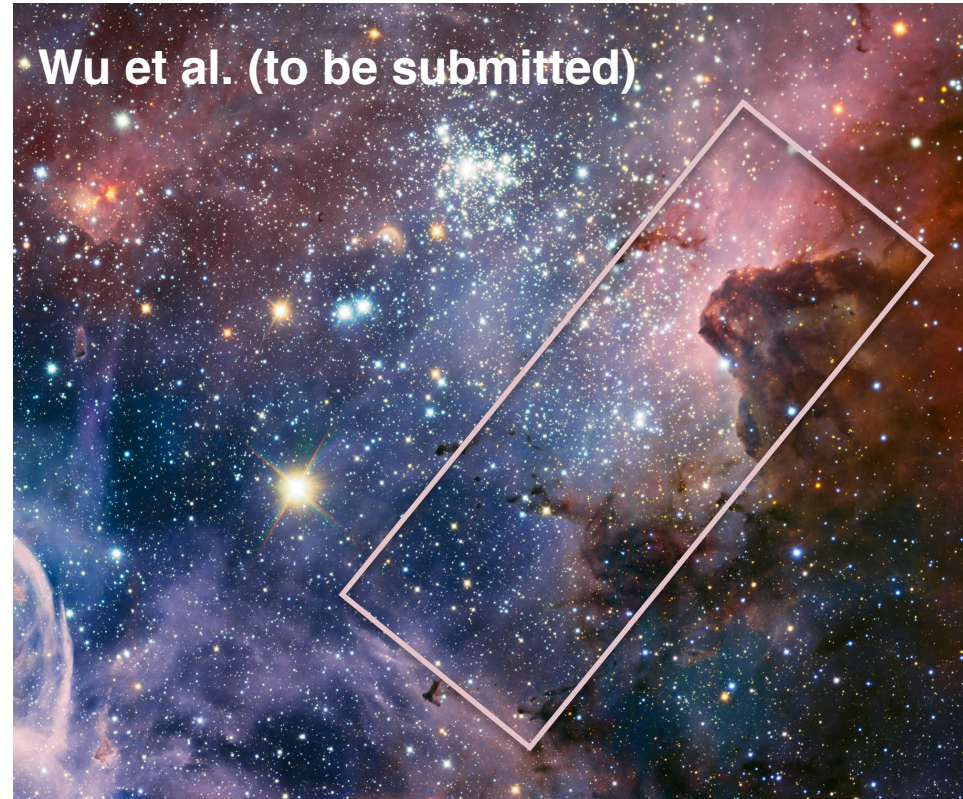
Example of IRAM observations

#### **Orion B project (Pety et al. 2017):**

- Large program IRAM-30m/EMIR
- 141 050 pixels
- tens of line intensities at each pixel

→ need new techniques to compare  
models & observations





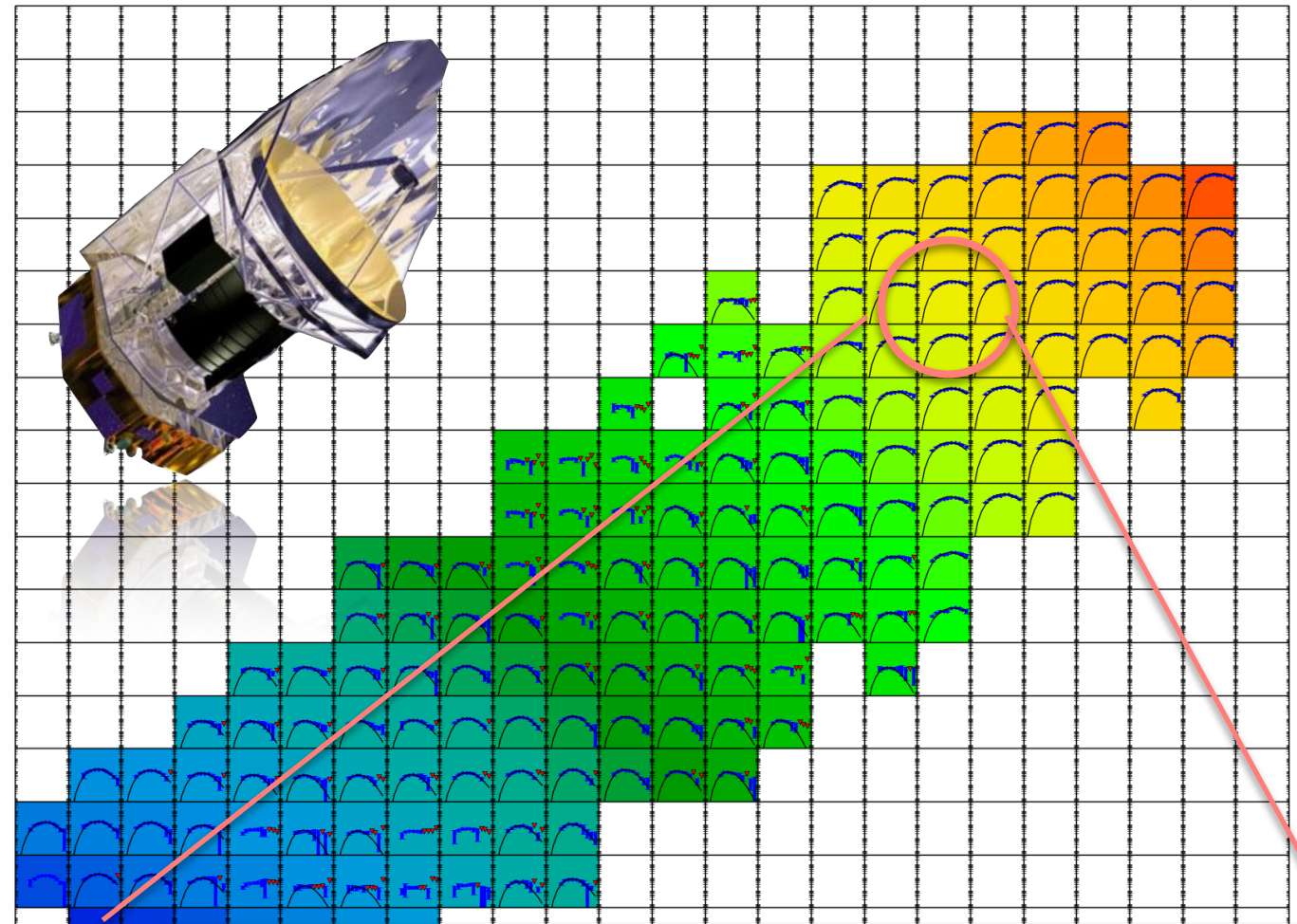
-59d  
39.0m  
12.12s

-59d  
37.0m  
42.24s

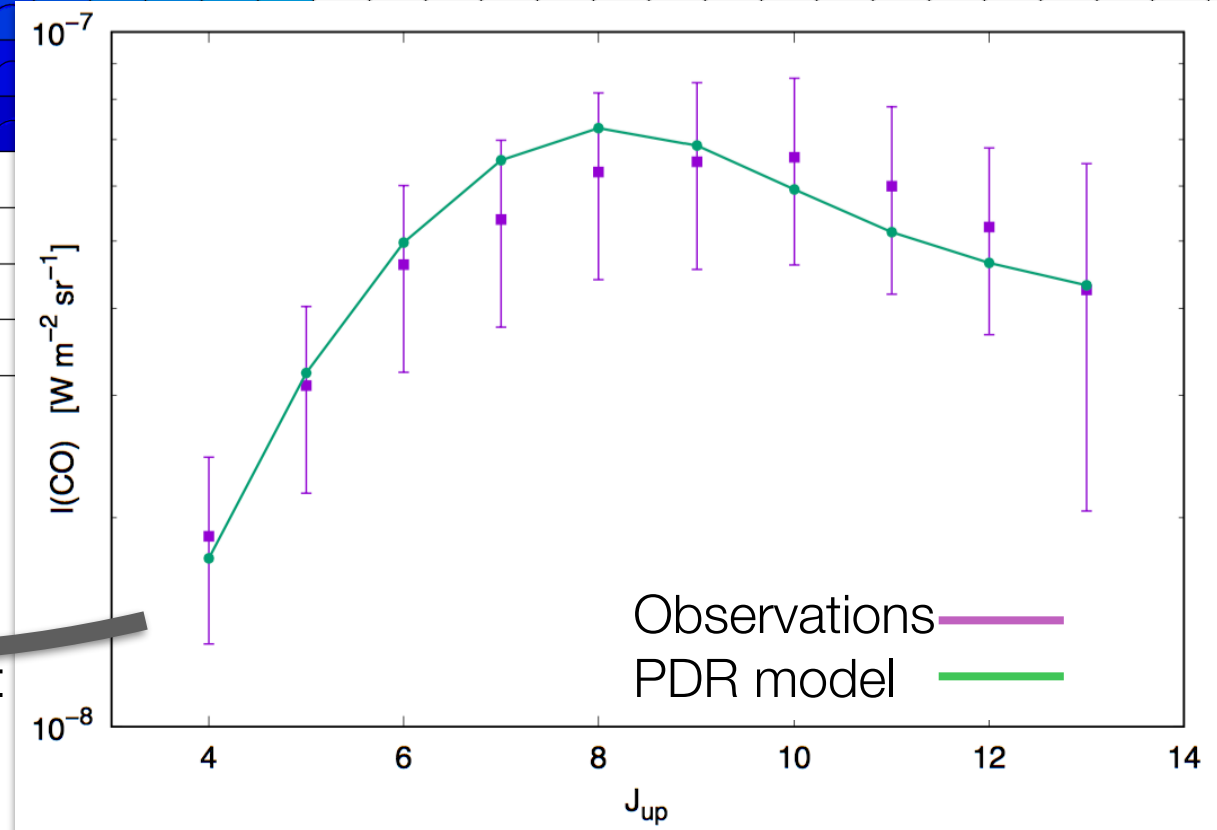
-59d  
36.0m  
12.29s

-59d  
34.0m  
42.28s

-59d  
33.0m  
12.21s  
10h  
44.0m  
7.84s

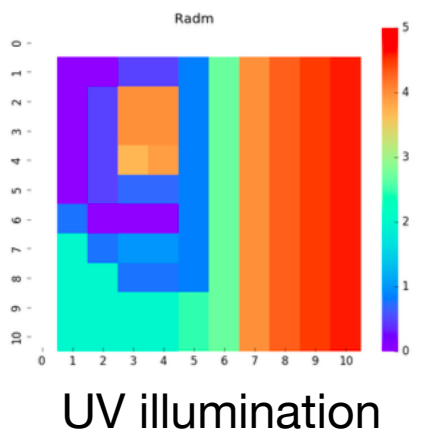
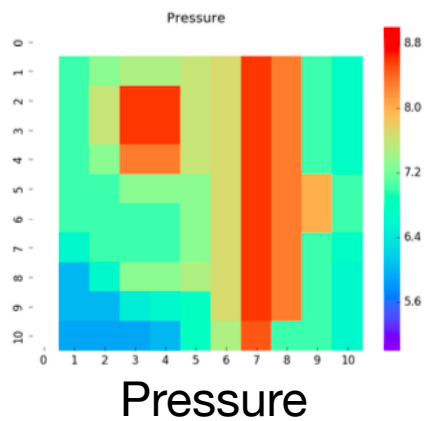


Find best  
physical  
parameters at  
each pixel

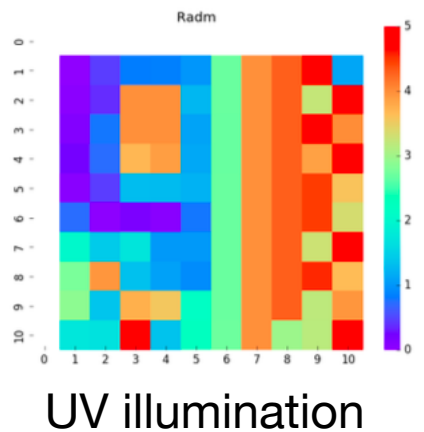
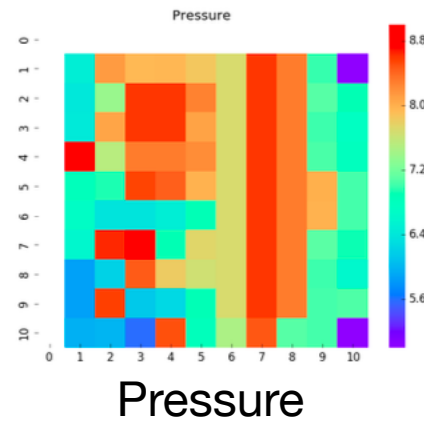


**Problem:** in low SNR regions, data insufficient to constrain the model.

True physical maps :



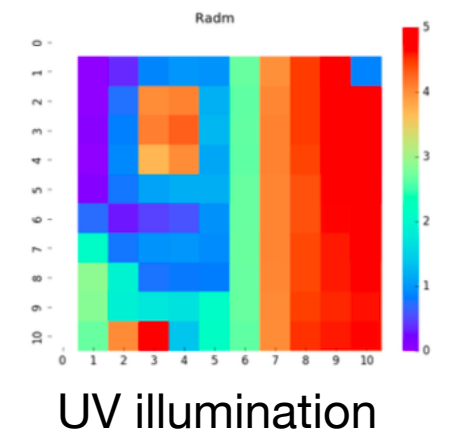
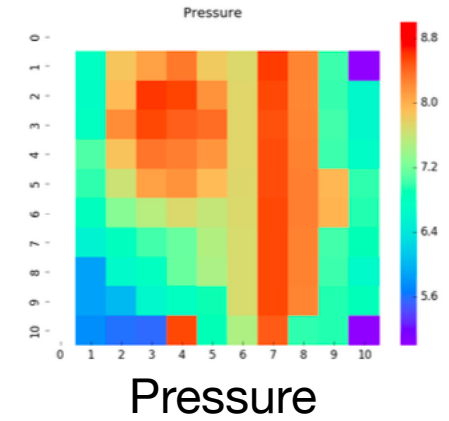
Pixel-by-pixel fit results :



## Solution: Regularization

- penalize non-smoothness in the map
- equivalent to bayesian fitting with a smoothness prior
- results in adaptive smoothing in order to have a well constrained fit

Regularized fit results :



(Based on noisy synthetic observations)

Work in progress.

Internship of Nicolas Chabaliar

# Example of what we would like

Best fit

click on pixels to view detailed info



CODES &amp; DATABASES

TECHNOLOGIES

PARTNERS

help

## ISMDB Map Fit

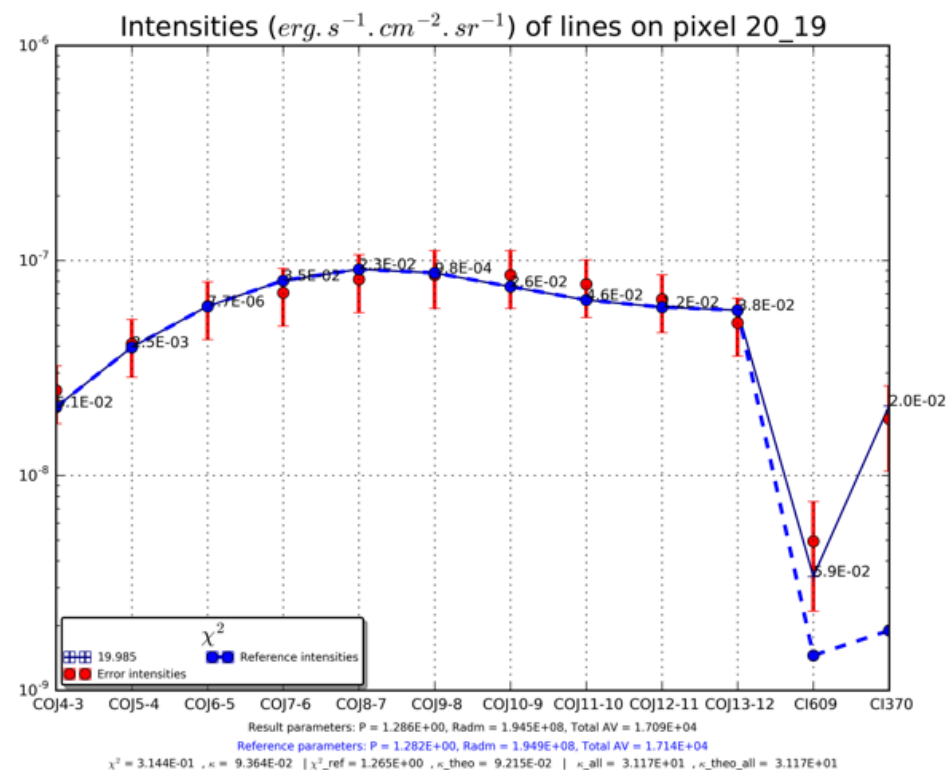
Fit map to PDR conditions

Choisir le fichier aucun fichier sél.

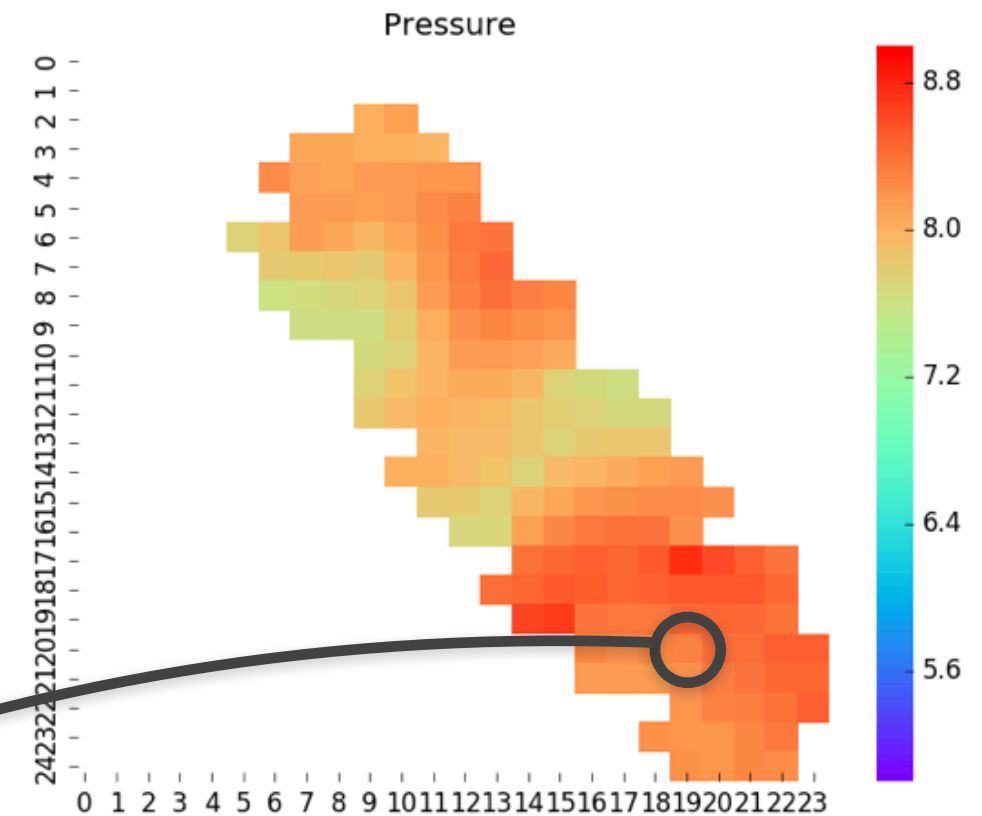
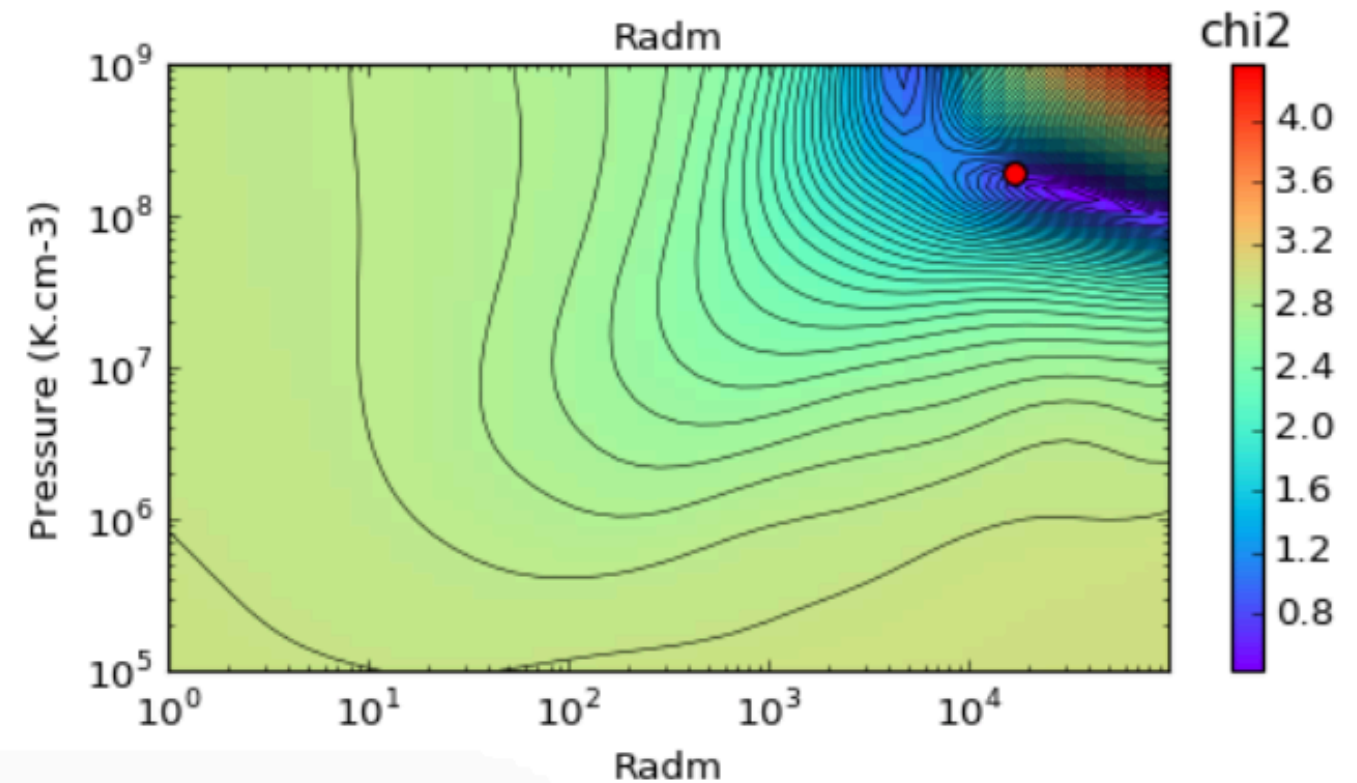
upload

### Fit detail for pixel 20,19

Best model fit detail



Fit solutions domain



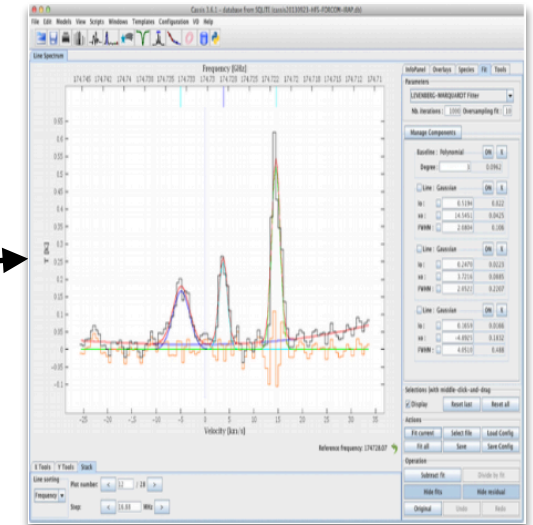


## ④ Coupling ISMDB with other services

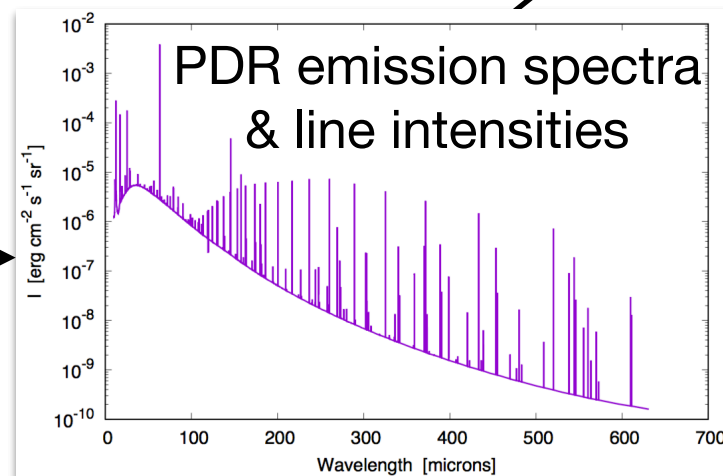
PDR spectra asked by several other services

- CASSIS (ANO5 - IRAP)
- ARTEMIX (ANO3 - ALMA Regional Center)
- CIGALE (ANO5 - GAZPAR)

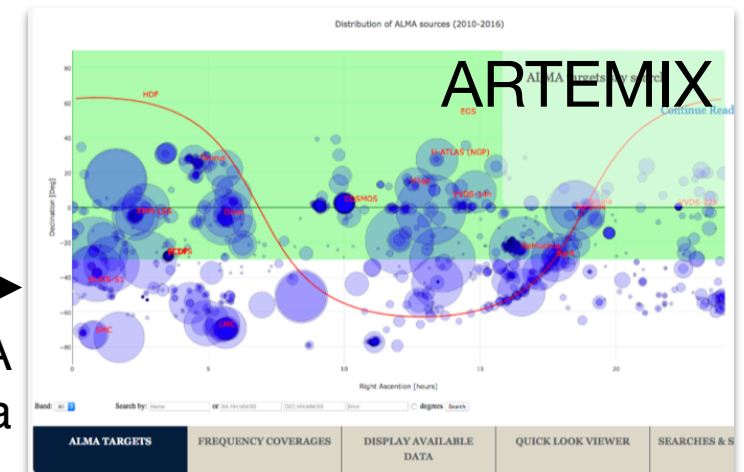
CASSIS



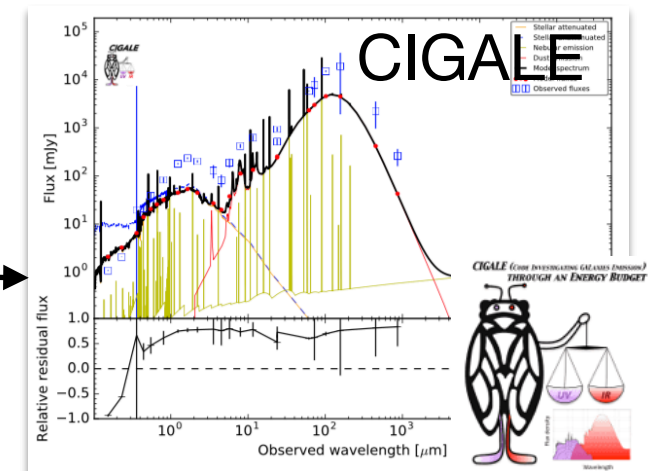
Interpret ISM spectra



Interpret ALMA archive data



Analyse galaxies molecular emission



CIGALE

## Major new developments:

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- ④ Links with other services